Accreting X-ray Pulsars with INTEGRAL

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Integral Workshop, Oct. 18., Sardinia

Accreting X-ray Pulsars with INTEGRAL

Vela X-I and 4U 1909+07

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INTEGRAL Science Data Centre

Vela X-I

- spin period: 283.5s
- normal: 200mCrab
- flares: up to 6 Crab
- 2 Msec public data











Vela X-I

no JEM-X data Vela X-1 a) 10^{0} normalized counts/sec/keV normalized counts/sec/keV 10_{-3} spectral model: Powerlaw x FDCut $I_{\rm cont}(E) \propto E^{-\Gamma} \times \frac{E - E_{\text{cut}}}{E_{\text{fold}}} + 1$ 10^{-5} 10 -b) X • CRSF: -1053.2 keV σ=7.5keV C) 2 × $\tau = 1.0$ 50 100 20 Channel Energy [keV]

50

Channel Energy [keV]

20

100

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Vela X-I

no JEM-X data Vela X–1 a) 10^{0} normalized counts/sec/keV normalized counts/sec/keV 10_{-3} spectral model: Powerlaw x FDCut no line! $I_{\rm cont}(E) \propto E^{-\Gamma} \times \frac{-}{\exp\left(\frac{E-E_{\rm cut}}{E_{\rm fold}}\right)+1}$ 10^{-5} 10 b) X • CRSF: -1053.2 keV σ =7.5keV F C) 2 × $\tau = 1.0$

Vela X-I

- Schanne et al.:
 3 Msec private data
- JEM-X data!
- model: cutoffpl
- two CRSFs!
- $E_1=27$ keV $\sigma=12$ keV $E_2=54$ keV $\sigma=11$ keV
- seen by ISGRI and SPI



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~25keV line is real!

Vela X-I: flares

- spectra from flares only
- CRSFs found at consistent energies



Vela X-I: flares

- spectra from flares only
- CRSFs found at consistent energies



 \Rightarrow no change due to luminosity

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4U 1909+07



4U 1909+07

data and folded model Model: 4U1909_sum_pha.fits 0.1 normalized counts/sec/keV cutoffpl 0.01 210-³ T \approx 0 Т 2 50 100

channel energy (keV)

TT

4U 1909+07



4U 1909+07



4U 1909+07

Ingo Kreykenbohm (IAAT)



20-40 keV 4 Rev = 0067 - 0070 P = 604.75 +/- 0.01 \overrightarrow{P} = 604.75 +/- 0.01

40 1909+07



Ingo Kreykenbohm (IAAT)